

6 Feb Finish Review & Start Symmetries

A.Y.A.C.I

Give ~15 min to work on packet followed by a quick debrief

Time

Goal & Objectives

By the end of this class, you should be able to...

- Describe the differences between simple surface-level symmetries & underlying symmetries in the equations of motion.
- Recognize that patterns and degeneracies are often the result of an underlying symmetry.
- Identify a symmetry underlying a pattern or degeneracy.
- Given a Lagrangian w/ a symmetry, determine the corresponding conserved quantity.
- State Noether's Theorem and give a few examples.

Where I am pulling from

For this discussion of symmetries and the related topic of group theory, I am pulling from several places

- Griffiths, Chapter 6
 - Griffiths Intro to Elementary Particles, Ch 4
 - Matthews & Walker 2nd Ed., Ch 16
 - Marion & Thornton Classical Dynamics, Ch 7
- } etc

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Motivation: word cloud

Reviewing Mathematical Symmetries

Show $f(x) = e^{-x^2} \sin(x^3)$

- Give me a symmetry

-) It's odd

-) Animate

- How are the following related?

-) $[f(-x)]^6$? $[f(x)]^6 =$

-) $\left. \frac{\partial f}{\partial x} \right|_{+2}$? $\left. \frac{\partial f}{\partial x} \right|_{-2} =$

-) $\int_{-3}^3 f(x) dx = 0$

From the symmetry of the function, we have learned a lot!

Introduction to Symmetries in Physics

Physics is often the reverse:

- 1) We find a neat pattern:

- a) The 1st excited states of a particle in a 2-D box are two-fold degenerate:

$n_x=1, n_y=2$ vs. $n_x=2, n_y=1$

- b) Planets return to the same spot each "year"

- 2) These patterns are then found to be due to some underlying symmetry

- a) The 2-D box is the same if I rotate by 90°

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E.O.M.: Equation of Motion.

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b) Newton's Law of Gravitation
(& GR for that matter)

$$F = -\frac{GMm}{r^2}$$

is spherically symmetric: Invariant under rotations

This planetary orbit example is particularly important as it shows us the types of symmetries most important for physics. Slide

- The ancient Greeks thought that the planets must move in circles as that is the most symmetric shape
 -) Obviously wrong: ellipses
- Newton discovered that, while the paths themselves are not spherically symmetric, the underlying E.O.M. are:

$$m\ddot{r} = -\frac{GMm}{r^2}$$

Interesting Aside: Did Newton use the fact that the gravitational force "ought" to be spherically symmetric (effectively following Aristotle & co.) to figure out GMm/r^2 ? What if I hadn't been?

w.r.t.: with respect to

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Going Deeper: Some "Dumb" Patterns

Let's do some "dumb" experiments:

- Drop an iolab & time the fall
 - > Different times
 - > Different places
 - > Different orientations
- } Same Results!

Just like how the planets return to their original positions, we notice a symmetry

- The laws of physics seem invariant w.r.t. translations in time
- & space
- & orientation

Are these symmetries hinting at something deeper like how the closed orbits of planets ~~bequite~~ indicate the spherical symmetry of GMm/r^2 ?

Let's look at spatial invariance

Eg

Consider a Lagrangian in rectangular coordinates

$$L(x_i, \dot{x}_i, t)$$

(Following MB&T 7.9 here)

We will only consider

- Closed systems $\partial L / \partial t = 0$
- Speed-independent potentials

$$\partial L / \partial \dot{x}_i = 0 \quad \forall \dot{x}_i$$

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$\therefore \rightarrow$ Therefore

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E.g. the iOLab's Lagrangian

$$L = \frac{1}{2} m \dot{y}^2 - mgy$$

Let's shift our entire system an infinitesimal amount:

$$x_i \rightarrow x_i + \delta x_i$$

The change in L , δL , is (chain rule)

$$\delta L = \frac{\partial L}{\partial x_i} \delta x_i + \frac{\partial L}{\partial \dot{x}_i} \delta \dot{x}_i$$

(Einstein convention of sum over i implied)

What is $\delta \dot{x}_i$?

Zero: all we did was slide our origin which does not change $\dot{x} = \frac{dx}{dt}$

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$$\therefore \delta L = \frac{\partial L}{\partial x_i} \delta x_i$$

Recall the principle of least action:

- $S \equiv \int L dt \leftarrow$ Action
- Particles follow the path of least action

\rightarrow Minimizing $S \rightarrow \delta S = 0 \rightarrow \delta L = 0$

\rightarrow This is what you use to derive the Lagrangian e.o.m. $\frac{d}{dt} \frac{\partial L}{\partial \dot{x}_i} = \frac{\partial L}{\partial x_i}$

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$\therefore \rightarrow$ because

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In order for $\delta L = 0$,

$$\frac{\partial L}{\partial x_i} = 0 \quad \forall x_i$$

Coming back to the Euler-Lagrange E.O.M.

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{x}_i} = \frac{\partial L}{\partial x_i}$$

$$\frac{\partial L}{\partial x_i} = 0 \rightarrow \frac{\partial L}{\partial \dot{x}_i} = \text{constant}$$

Now, putting in $L = T - U$ recalling

- $T = \frac{1}{2} m \dot{x}_i^2$
- $\partial U / \partial \dot{x}_i = 0$

$$\frac{\partial L}{\partial \dot{x}_i} = \text{constant} \quad 0 \because \partial U / \partial \dot{x}_i = 0$$

$$\frac{\partial}{\partial \dot{x}_i} (T - U) = \text{constant}$$

$$\frac{\partial}{\partial \dot{x}_i} \left(\frac{1}{2} m \dot{x}_i^2 \right) = \text{constant}$$

$$\sum_i m \dot{x}_i = \text{constant} \leftarrow \text{Momentum is conserved}$$

\therefore A translational symmetry in space \rightarrow ~~***~~
momentum conservation ~~***~~

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Noether's Theorem

Generalizes what we did!

<u>Symmetry</u>	<u>Conserved Quantity</u>	<u>Slide</u>
Space Translation	Momentum	
Time Translation	Energy	
Orientation / Rotation	Angular Momentum (& Boosts)	
EM Gauge $\phi \rightarrow \phi + \lambda(x)$	Electric Charge	

Einstein's Obit

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