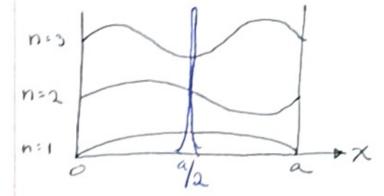


Homework 7

1. (Griffiths 7.1) Suppose we put a delta-for bump at the center of the square well: $H' = \chi \, \delta(\chi - \sqrt[4]{a})$

2) Find the first order perturbed energies. Why are energies not perturbed for even

The second port is straightforward: all even-n are zero at the center of the well so are unaffected by the $\delta(\chi-\frac{7}{2})$.



Now to find the energy corrections:

$$E_{n}^{(i)} = \langle n | H' | n \rangle$$

$$= \int dx \sin \left(\frac{n\pi x}{a} \right) d \delta \left(\chi - \frac{\pi}{a} \right) \sin \left(\frac{n\pi x}{a} \right) \cdot \left(\frac{\pi}{a} \right)^{2}$$

$$= \frac{2}{a} \propto 8 \sin^{2} \left(\frac{\pi n}{a} \cdot \frac{x}{a} \right) = \frac{2}{a} \sin^{2} \left(\frac{n\pi}{a} \right)$$

12

b) Find the first 3 non-zero terms in

 $\Psi_{n}^{(1)} = \sum_{m \neq n} \frac{\langle \psi_{m}^{(0)} | H' | \psi_{m}^{(0)} \rangle}{E_{n}^{(0)} - E_{m}^{(0)}} | \psi_{m}^{(0)} \rangle$

with nol

Y(1) = \(\frac{\partial \partial \part

Let's look at the numerator:

 $\langle \Psi_{m}^{(0)} | H' | \Psi_{n}^{(0)} \rangle = \int dx \left(\frac{1}{a} \right)^{2} \sin \left(\frac{m\pi x}{a} \right) \alpha \delta(x - \frac{\alpha}{a}) \sin \left(\frac{\pi x}{a} \right)$

= $\frac{2d}{a} \sin\left(\frac{m\pi}{2}\right) \sin\left(\frac{\pi}{2}\right)$

Sin (T/2) = 1 so

(40) | H' | 400) = 2x sin (mit)

which is non-zero for odd n. Since m=1 is excluded, we have m = 23, 5, 7}

Meanwhile, in the denominator, we have

$$E_n^{(0)} - E_m^{(0)} = \frac{\pi t^2 n^2}{2ma^2} - \frac{\pi^2 t^2 m^2}{2ma^2}$$
, but n=1 so $= \frac{\pi^2 t^2}{2ma^2} \left(1 - m^2\right)$

and

Putting it all together, and recognizing that sin (311/2)=1, sin (511/2)=1, sin (511/2)=1, sin (711/2) 2-1, we have

9 HW#7

2A

Now perturb by k > (+ E)k.

a) Find the exact solution and expand to 2nd order in E.

So the energies are En= (n+ =) thw, or, in terms of k & E:

$$E_n = h(n+1/2)\sqrt{k(1+\epsilon)^2} = h\omega(n+1/2)\sqrt{1+\epsilon^2}$$

$$(1+\varepsilon)^n = 1 + n\varepsilon + \frac{1}{2}n(n-1)\varepsilon^2 + ...$$

$$E_{n} = (n \cdot 2) \hbar \omega \left[1 + \frac{1}{2} \epsilon + \frac{1}{2} \left(\frac{1}{2} \right) \left(-\frac{1}{2} \right) \epsilon^{2} + \dots \right]$$

b) Now use perturbation theory (2.1thout integrals!) to find En and see that it agrees with part a.

En = (n|H'|n)

What is H'?

H= p2 + k(1+E) x2. Nov expand:

 $H = \frac{p^2 + k x^2 + k \epsilon x^2}{2m \cdot 2}$

 $: E_n^{(i)} = \langle n | H' | n \rangle$ $= \langle n | \frac{k \epsilon}{2} x^2 | n \rangle$

] 2 ways to do this:

1) The virial theorem $\langle T \rangle = \frac{1}{2} \langle V \rangle \xrightarrow{\text{tor}} \langle T \rangle = \langle V \rangle$ as χ^n with n=22) Operators

The virial theorem method is most straightforward $E_n^{(1)} = \varepsilon \langle n | V | n \rangle = \varepsilon \langle V \rangle$

HW #7

2

Since (V)= (T) and (V)+ (T)= En,

En = = = = = (n+1/2) to

a perfect match for the linear term highlighted at the bottom of pg 2A.

The other option is to write x: to (a+a) and proceed that way.

Problem 3 (Griffiths 7.4)

Start by stating global assumptions for this notebook: that both energies are real and positive while $0 < \lambda < 1$.

```
\label{eq:local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_local_
```

Apply perturbation theory to the most general two-state system with unperturbed energies:

```
In[2]:= MatrixForm[H0 = {{Ea, 0}, {0, Eb}}]

Out[2]//MatrixForm=

(Ea 0
0 Eb)
```

and perturbation

Out[5]//MatrixForm=

 $egin{pmatrix} \mathsf{Ea} + \mathsf{Vaa} \ \lambda & \mathsf{Vab} \ \lambda & \mathsf{Eb} + \mathsf{Vbb} \ \lambda \end{pmatrix}$

Simplifies to

$$\begin{aligned} & \text{In} [6] \coloneqq \text{ Emp = Eigenvalues [H]} \\ & \text{Out} [6] \coloneqq \begin{cases} \frac{1}{2} \left(\mathsf{Ea} + \mathsf{Eb} + \mathsf{Vaa} \, \lambda + \mathsf{Vbb} \, \lambda - \\ & \sqrt{\left(\left(- \mathsf{Ea} - \mathsf{Eb} - \mathsf{Vaa} \, \lambda - \mathsf{Vbb} \, \lambda \right)^2 - 4 \, \left(\mathsf{Ea} \, \mathsf{Eb} + \mathsf{Eb} \, \mathsf{Vaa} \, \lambda + \mathsf{Ea} \, \mathsf{Vbb} \, \lambda - \mathsf{Vab}^2 \, \lambda^2 + \mathsf{Vaa} \, \mathsf{Vbb} \, \lambda^2 \right) \right) \right), \\ & \frac{1}{2} \left(\mathsf{Ea} + \mathsf{Eb} + \mathsf{Vaa} \, \lambda + \mathsf{Vbb} \, \lambda + \\ & \sqrt{\left(\left(- \mathsf{Ea} - \mathsf{Eb} - \mathsf{Vaa} \, \lambda - \mathsf{Vbb} \, \lambda \right)^2 - 4 \, \left(\mathsf{Ea} \, \mathsf{Eb} + \mathsf{Eb} \, \mathsf{Vaa} \, \lambda + \mathsf{Ea} \, \mathsf{Vbb} \, \lambda - \mathsf{Vab}^2 \, \lambda^2 + \mathsf{Vaa} \, \mathsf{Vbb} \, \lambda^2 \right) \right) \right) \right)} \\ & \text{Which simplifies into E- (Em) and E+ (Ep) via} \\ & \text{In} [7] \coloneqq \text{Simplify} \left[\left(\mathbf{a} + \mathbf{b} \right) \, ^2 - 4 \, \mathbf{a} \, * \, \mathbf{b} \right] \\ & \text{Out} [7] = \left(\mathbf{a} - \mathbf{b} \right)^2 \\ & \text{In} [8] \coloneqq \text{Em = 1 / 2 } \times \left(\left(\mathsf{Ea} + \mathsf{Eb} + \mathsf{Vaa} \, * \, \lambda + \mathsf{Vbb} \, * \, \lambda \right) - \mathsf{Sqrt} \left[\left(\mathsf{Ea} + \lambda * \mathsf{Vaa} - \mathsf{Eb} - \lambda * \, \mathsf{Vbb} \right) \, ^2 + 4 \, * \, \lambda \, ^2 \, * \, \mathsf{Vab} \, ^2 \right] \right) \\ & \text{Out} [8] \coloneqq \frac{1}{2} \left(\mathsf{Ea} + \mathsf{Eb} + \mathsf{Vaa} \, \lambda + \mathsf{Vbb} \, \lambda - \sqrt{4 \, \mathsf{Vab}^2 \, \lambda^2 + \left(\mathsf{Ea} - \mathsf{Eb} + \mathsf{Vaa} \, \lambda - \mathsf{Vbb} \, \lambda \right)^2} \right) \\ & \text{In} [9] \coloneqq \text{Ep = 1 / 2} \times \left(\left(\mathsf{Ea} + \mathsf{Eb} + \mathsf{Vaa} \, * \, \lambda + \mathsf{Vbb} \, * \, \lambda \right) + \mathsf{Sqrt} \left[\left(\mathsf{Ea} + \lambda * \mathsf{Vaa} - \mathsf{Eb} - \lambda * \, \mathsf{Vbb} \right) \, ^2 + 4 \, * \, \lambda \, ^2 \, * \, \mathsf{Vab} \, ^2 \right] \right) \\ & \text{Out} [9] \coloneqq \frac{1}{2} \left(\mathsf{Ea} + \mathsf{Eb} + \mathsf{Vaa} \, \lambda + \mathsf{Vbb} \, \lambda + \sqrt{4 \, \mathsf{Vab}^2 \, \lambda^2 + \left(\mathsf{Ea} - \mathsf{Eb} + \mathsf{Vaa} \, \lambda - \mathsf{Vbb} \, \lambda \right)^2} \right) \end{aligned}$$

b) Expand your results from (a) to second order in λ . Show that the results are the same as what we obtained in perturbation theory.

Both of which match our results from perturbation theory!

c) Setting Vaa = Vbb = 0, show that the series only converges if | Vab / (Eb - Ea) | < 1/2

The series only converge if

In[14]:=

Out[14]=

$$\frac{4 \text{ Vab}^2}{\left(-\text{ Ea} + \text{ Eb}\right)^2} \, < \, 1$$

or equivalently

$$In[15]:= | Vab / (Eb - Ea) | < 1 / 2$$

HW#7

4A

$$H' = \frac{1}{4\pi\epsilon_0} \left(\frac{e^2}{R} - \frac{e^2}{(R-\chi_1)} - \frac{e^2}{(R+\chi_2)} + \frac{e^2}{(R-\chi_1+\chi_2)} \right)$$

te -e telection

(a) Explain the perturbation.

The first term is the nuclear repulsion: + 1/R

The second term is the attraction between the electron on the left and the nucleus on the right: -ez/(r-x,)

The third term is the attraction between the electron on the right and the left nucleus:

The final term is the repulsion of the $t \ge 0$ electrons: $t \in \mathbb{R}^2$ $\mathbb{R}^2 - x_1 + x_2$

Also, assuming $|\chi_1| \ll R \in |\chi_2| \ll R$, show $|H' \simeq -\frac{e^2 \chi_1 \chi_2}{2\pi \epsilon_1 R^3}$

I appears in all terms but the first in R-X some form. We want (1+E)

$$\frac{1}{R-\chi} = \frac{1}{R} \left(\left| -\frac{\chi}{R} \right|^{-1} \right)$$

$$\sim \frac{1}{R} \left(\left| -\frac{\chi}{R} + \frac{1}{2} \left(- \right) \left(- \right| - 1 \right) \left(\frac{\chi}{R} \right)^2 + \cdots \right)$$

$$\sim \frac{1}{R} \left(\left| -\frac{x}{R} + \left(\frac{x}{R} \right)^2 + \cdots \right) \right)$$

$$H' \approx \frac{e^{2}}{4\pi\epsilon_{0}R} \left[\sqrt{-\left(\sqrt{\frac{x_{1}}{R}} + \frac{x_{1}^{2}}{R^{2}} \right) - \left(\sqrt{-\frac{x_{2}}{R}} + \frac{x_{2}^{2}}{R^{2}} \right) + \left(\sqrt{\frac{x_{1} - x_{2}}{R}} + \left(\frac{x_{1} - x_{2}}{R^{2}} \right) - \left(\sqrt{-\frac{x_{2}}{R}} + \frac{x_{2}^{2}}{R^{2}} \right) \right] + \left(\sqrt{\frac{x_{1} - x_{2}}{R}} + \left(\frac{x_{1} - x_{2}}{R^{2}} \right) - \left(\sqrt{-\frac{x_{2}}{R}} + \frac{x_{2}^{2}}{R^{2}} \right) \right]$$

$$\approx \frac{e^2}{4\pi\epsilon_0 R} \left[-\frac{\chi}{R} + \frac{\chi}{R} + \left(\frac{\chi}{R} - \frac{\chi^2}{R} \right) - \frac{\chi^2}{R} + \frac{\chi^2}$$

$$\approx -\frac{e^2 \chi_1 \chi_2}{2 \pi \epsilon_0 R^2} \quad 0. \epsilon. \delta.$$

14

HW #7

40

5) Show that Ho plus the simplified H' separates into 2 QMHO's:

$$H = \left[\frac{p_{1}^{2}}{2\pi} + \frac{1}{2} \left(k - \frac{e^{2}}{2\pi \epsilon_{0} R^{3}} \right) \chi_{+}^{2} \right] + \left[\frac{p_{2}^{2}}{2\pi} + \frac{1}{2} \left(k + \frac{e^{2}}{2\pi \epsilon_{0} R^{3}} \right) \chi_{-}^{2} \right]$$

with $\chi_t = \frac{1}{\sqrt{2}} \left(\chi_1 = \chi_2 \right)$ $p_t = \frac{1}{\sqrt{2}} \left(p_i = p_2 \right)$

The casiest wary to do this is to work backwards and expand the result

(since 162 are independent particles, P, & Pz commute!)

Similarly:

c) Assuming
$$k \gg \left(\frac{e^2}{2\pi\epsilon_0 R^3}\right)$$
 Show that $\Delta V = E - E_0 = -\frac{\hbar}{8m^2\omega_0^3} \left(\frac{e^2}{2\pi\epsilon_0}\right)^2 \frac{1}{R^6}$

The energies, in this case, are

$$\Delta V \approx \frac{1}{2} \hbar \omega_0 \left[\left[-\frac{1}{2} \left(\frac{e^2}{2\pi \epsilon_0 R^3 k} - \frac{1}{8} \left(\frac{e^2}{2\pi \epsilon_0 R^3 k} \right)^2 \right] + \left[\left[+\frac{1}{2} \left(\frac{e^2}{2\pi \epsilon_0 R^3 k} - \frac{1}{8} \left(\frac{e^2}{2\pi \epsilon_0 R^3 k} \right)^2 \right] \right] - \hbar \omega_0$$

$$\Delta V \approx \frac{1}{2} \hbar \omega_0 \left[\left[\frac{e^2}{2\pi \epsilon_0 R^3 k} - \frac{1}{8} \left(\frac{e^2}{2\pi \epsilon_0 R^3 k} \right)^2 \right] - \hbar \omega_0$$

At first order, things go to zero

$$= -\frac{e^2}{2\pi \epsilon_0 R^3} \left\langle 00 \mid \chi_1 \chi_2 \mid 00 \right\rangle$$

Since $\chi_1 = \frac{1}{2} \left(\alpha_1 + \alpha_1^+ \right)$, we see that the result is zero: α_1 by lowering $|0\rangle \Rightarrow 0$ and α_1^+ by orthogonality. The same will be true for χ_2

At second order

$$E_{0}^{(2)} = \sum_{n_{1}=1}^{\infty} \sum_{n_{1}=1}^{\infty} \frac{|KOO|H'|n_{1}n_{2}|^{2}}{|E_{00}-E_{n_{1}}n_{2}|}$$

$$= \left(\frac{e^{2}}{2\pi\epsilon_{0}R^{3}}\right)^{2} \sum_{n_{1}=1}^{\infty} \sum_{n_{2}=1}^{\infty} \frac{1}{|E_{00}-E_{n_{1}}n_{2}|} \left|\langle OO|\chi_{1}\chi_{2}|n_{1}n_{2}\rangle\right|^{2}$$

Again, recalling $x_i = \sqrt{\frac{1}{2}} \left(\alpha_i + \alpha_i^{\dagger}\right)$ the only non-zero term is III) as all others go to zero by orthogonality.

$$\left| \left\langle 00 \left| \chi_1 \chi_2 \right| 11 \right\rangle \right|^2 = \left| \sqrt{\frac{\hbar}{2m\omega}} \sqrt{\frac{\hbar}{2m\omega}} \sqrt{1} \sqrt{1} \right|^2$$

$$=\left(\frac{1}{2m\omega}\right)^{2}$$

$$F_0^{(2)} = -\frac{1}{8m^2\omega^3} \left(\frac{e^2}{2\pi\epsilon_0}\right)^2 \frac{1}{R^6} \quad 0. \, \epsilon. \, \delta.$$